MAGIC Observation of the Galactic Center

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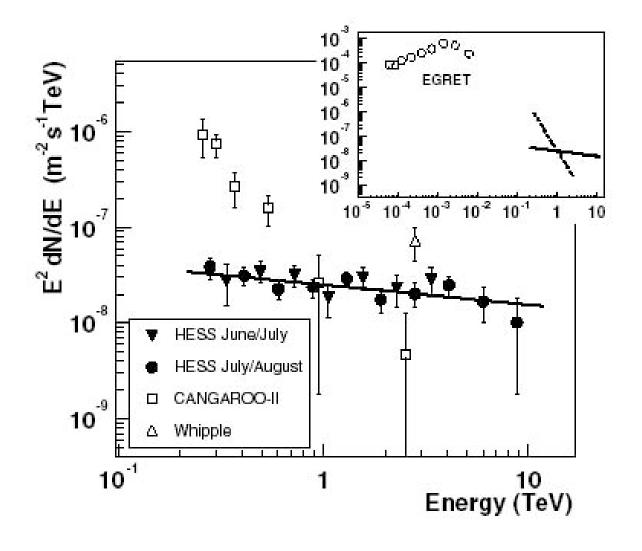
(most of the material taken from H. Bartko et al. for MAGIC, ICRC2005)

Outline

- introduction
- observations under large zenith angle
- source location
- first spectrum

GC observed in VHE gamma by Cangaroo, HESS, Whipple

· HESS and Cangaroo results don't match



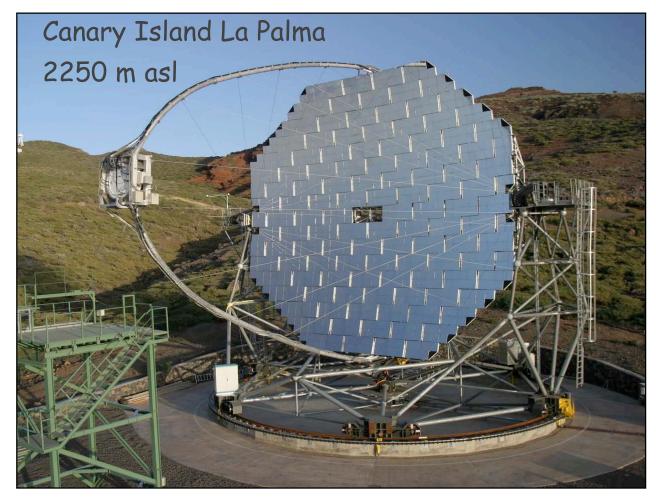
Cangaroo spectral index

_=-4.6±0.5

HESS spectral index

_=-2.63±0.04

The MAGIC Telescope



Largest Imaging Air Cherenkov Telescope

for γ -ray astronomy:

- 17 m ⊗mirror dish
- 3.5° field of view
- enhanced QE PMTsdesign aimed for:
- low energy threshold E_{γ} > 30 GeV
- fast repositioning
 t_R<30 s

In final configuration since September 2004 [NIM A518 (2004) 188]

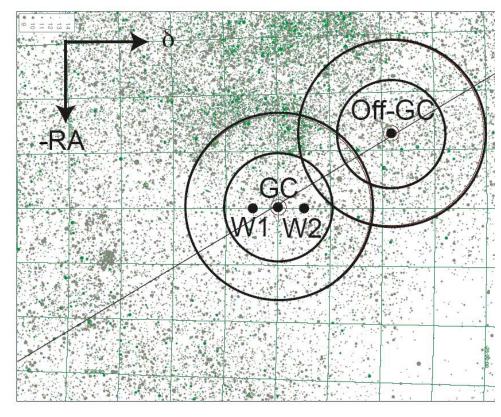
Observations

GC: culminates at ~58 deg ZA in La Palma few very bright stars, but inhomogeneous star field

-> dedicated OFF observations or wobble mode observations

MAGIC data:

- initial observ. in Sept. 2004,
 ~2.5 h only ON data,
 60°<= ZA<=70°
 not included in this analysis
- May 2005: ~7 h wobble data,
 58°<= ZA<=62°
- June/July 2005, ~15h ON,
 ~12h OFF data,
 58°<= ZA<=62°



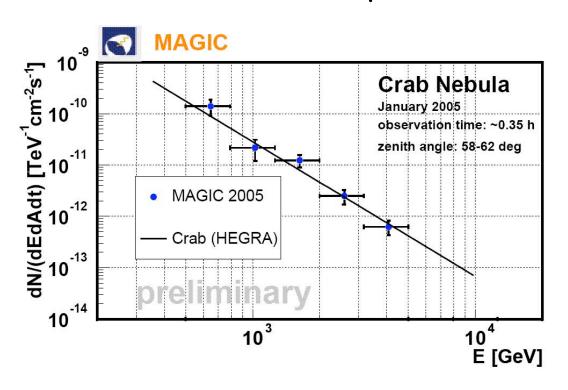
Analysis Chain

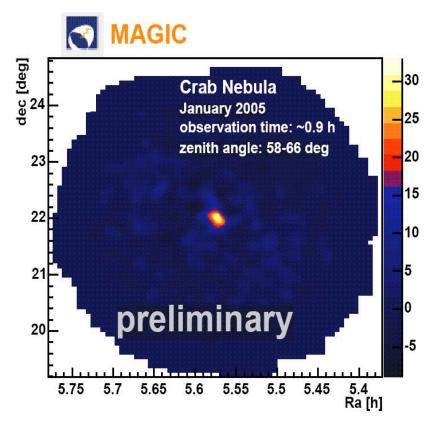
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data + MC:
    calibration
    absolute image cleaning: 10/5 photo electrons threshold
    calculation of Hillas image parameters
    Random Forest algorithm for gamma/hadron separation,
    energy estimation [NIM A516 (2004) 511]
    RF training: MC gamma - GC OFF data
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- -> gamma signal (ALPHA, DISP analysis)
- -> sky-maps, spectra

Test: Crab at Large Zenith Angle

- to test the analysis chain at large ZA: analyzed Crab data
- due to large ZA (\sim 60°): higher threshold (\sim 700 GeV), but also larger effective collection area -> good sensitivity
- flux: consistent with previous measurements (HEGRA)



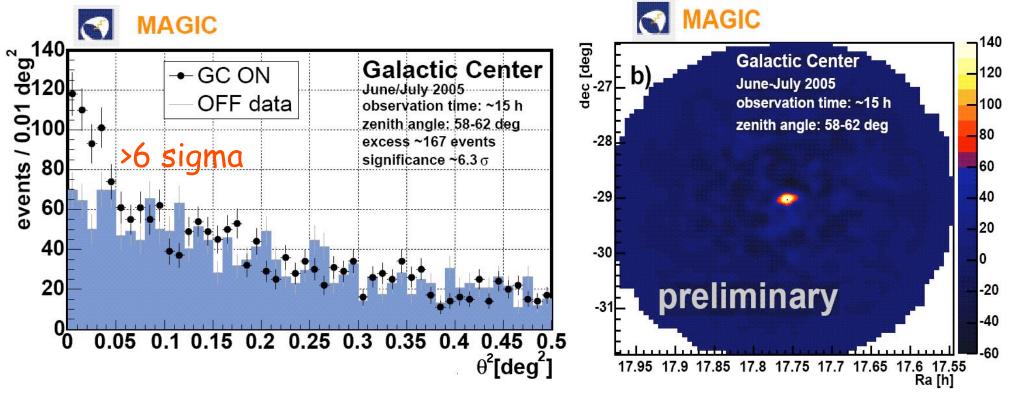


GC: Sky Maps

GC ON/OFF data: 15h/8h, June+July 2005, more data on tape

excess location: (Ra,Dec)=(17^h45'20", -29°2'), ±0.05° sys. error

· Galactic Center: within resolution compatible with point source

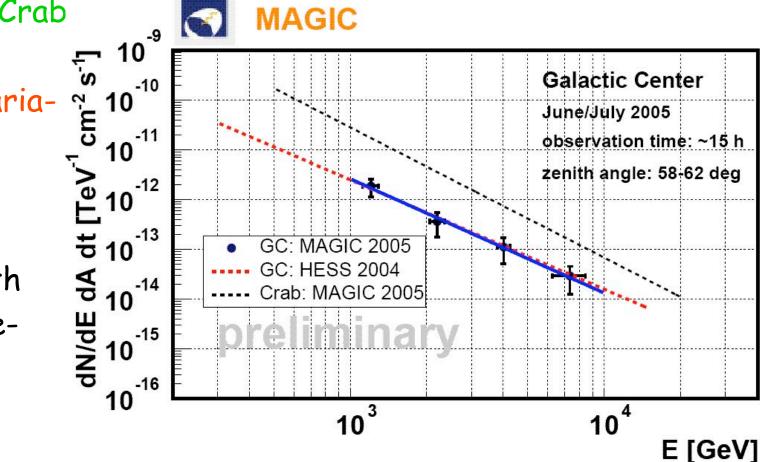


GC: Preliminary Spectrum

rather hard spectrum, _=-2.3±0.4

flux: ~10% of Crab
 above 1 TeV

- above 1 TeV
 no apparent variability during
 2 month of observations
- compatible with HESS measurements, not w/ Cangaroo



Conclusions

- Galactic Center has been observed by MAGIC under large ZA
- excess compatible with point-like source
- no apparent variability during 2 month of observations in 2005
- relatively hard spectrum, about 0.1 Crab above 1 TeV, compatible with HESS results
- possible source for inter-calibration of different IACTs